The geodesic structure of some spacetimes with symmetries — local and global properties

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Motivation: 'twin paradox'

The 'twin paradox' in SR has 3 levels of comprehending:

- why there is at all the asymmetry between the twins,
- why the accelerated twin is younger ('reverse triangle inequality'),
- what happens in a curved spacetime.
- Curvature \Rightarrow multitude od diverse results. Purely geometrical problem : which timelike curve joining two given points is the longest one? There is no *shortest* timelike line.
- The problem deeply enters into the geodesic structure of the spacetime this is why it is worth studying.
- The problem: local and global. Here: only a brief introduction.

Locally maximal timelike curves

A bundle of nearby timelike curves emanating from p and intersecting at p_1 , we seek for the longest one. It contains a geodesic γ_0 , $\gamma_0(0) = p$, $\gamma_0(s_1) = p_1$. The bundle may contain other geodesics γ_{ε} infinitesim. close to γ_0 , $|\varepsilon| \ll 1$. $x^{\mu}(s)$ — coordinates of γ_0 , $u^{\mu}(s)$ — tangent to γ_0 , $\bar{x}^{\mu}(s,\varepsilon)$ - coordinates of γ_{ε} ,

$$\bar{x}^{\mu}(s,\varepsilon) = x^{\mu}(s) + \varepsilon Z^{\mu}(s) + \delta^2 x^{\mu}(s) + \ldots + \delta^n x^{\mu}(s) + \ldots,$$

 $Z^{\mu}(s)$ — the connecting vector in the linear approximation, (the geodesic deviation vector, Jacobi vector field), $Z^{\mu} u_{\mu} = 0$, the *n*-th deviation $\delta^{n} x^{\mu} = O(\varepsilon^{n})$ for n > 1 is not a vector.

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 $Z^{\mu}(s)$ satisfies the geodesic deviation equation (GDE)

$$\frac{D^2}{ds^2} Z^{\mu} = R^{\mu}{}_{\alpha\beta\gamma} u^{\alpha} u^{\beta} Z^{\gamma}$$

(derived in the linear approximation). If $Z^{\mu} \neq 0$ and $Z^{\mu}(0) = Z^{\mu}(s_0) = 0$ for $0 < s_0 < s_1 \Rightarrow$ — either γ_{ε} intersects γ_0 at $q = \gamma_0(s_0)$ or — γ_{ε} is closer to γ_0 at q than at ε -distance. In both cases it is interpreted that γ_{ε} intersects γ_0 at q on the segment pp_1 . $q = \gamma_0(s_0)$ — point conjugate to p on γ_0 . Proposition (Hawking & Ellis)

A timelike geodesic γ_0 has the *locally maximal* length from p to p_1 iff there is NO point conjugate to p on the segment pp_1 .

Conversely:

if there is a point q conjugate to p on the segment pp_1 of γ_0 then there exists a nearby timelike curve λ (not necessarily geodesic) from p to p_1 which is longer than γ_0 , $s(\lambda) > s(\gamma_0)$.

The locally longest curve (geodesic) always exists in globally hyperbolic spacetimes.

CAdS — exist pairs of points connected by timelike lines but none is a geodesic and none is maximal (neither locally nor globally).

Criterion (Hawking & Ellis) If $R_{\alpha\beta} u^{\alpha} u^{\beta} \ge 0$ (SEC) on a timelike geodesic γ and if the tidal force $R_{\mu\alpha\nu\beta} u^{\alpha} u^{\beta} \ne 0$ at some point p_0 on γ_0 , then there is a pair of conjugate points p and q on γ_0 (if sufficiently extended). LOCALLY LONGEST GEODESICS We seek for maximal segments of geodesics \Rightarrow seek for conjugate points

 \Rightarrow seek for zeros of Jacobi fields $Z^{\mu}(s)$.

 γ_0 — chosen timelike geodesic, u^{μ} — tangent to γ_0 . Replace D^2/ds^2 in GDE by d^2/ds^2 . Choose tetrad $\{e_A^{\mu}\}$, A = 0, 1, 2, 3, orthonormal and parallelly transported along γ_0 ,

$$e_0^{\mu} \equiv u^{\mu}, \quad e_A^{\mu} e_{B\mu} = \eta_{AB} = \text{diag}(1, -1, -1, -1), \quad \frac{D}{ds} e_A^{\mu} = 0,$$

 $\{e_{a}{}^{\mu}\}$, a=1,2,3 —spacelike triad orthogonal to $\gamma_{0}.$ Expand

$$Z^{\mu} = \sum_{a=1}^{3} Z_a e_a{}^{\mu}, \quad \text{then GDE}$$

$$\frac{d^2}{ds^2} Z_a = -e^{\mu}_a R_{\mu\alpha\beta\gamma} u^{\alpha} u^{\beta} \sum_{b=1}^3 Z_b e_b{}^{\gamma},$$

Z_a — 3 Jacobi scalars.

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Killing vector $K^{\mu} \Rightarrow$ first integral of GDE

$$K_{\mu} \frac{D}{ds} Z^{\mu} - Z^{\mu} \frac{D}{ds} K_{\mu} = \text{const.}$$

PROCEDURE

- Choose a spacetime with some some symmetries (Killing vectors).
- Choose geometrically interesting timelike geodesic with explicit

$$x^{\alpha} = x^{\alpha}(\tau)$$
, e. g. $\tau = s$.

- Choose the spacelike triad e_a^{μ} as above.
- Solve GDE

$$rac{d^2}{ds^2} \, Z_a = - e^\mu_a \, R_{\mulphaeta\gamma} \, u^lpha \, u^eta \sum_{b=1}^3 Z_b \, e_b{}^\gamma$$

applying the first integrals and find a generic solution $Z_a(\tau)$.

— Consider all possible special solutions with $Z_a(\tau = 0) = 0$ and seek for their zeros,

$$Z_a(\tau_0)=0, \qquad \tau_0>0.$$

Then the geodesic $x^{\alpha} = x^{\alpha}(\tau)$ is *locally maximal* on the segment $0 \leq \tau < \tau_0$.

This is a *fully algorithmic* procedure for checking that the given geodesic is the unique *locally longest* curve between given points.

COMPUTATIONS

— Static spherically symmetric (SSS) spacetimes: de Sitter, CAdS, Schwarzschild, Reissner–Nordström and Bertotti–Robinson,

- ultrastatic SSS spacetimes: Barriola-Vilenkin monopole,
- cosmological Robertson-Walker spacetimes.
- Geodesics: radial and circular (if exist).
- A variety of diverse results. Spacetimes with similar symetries may have different geodesic structure: dS and CAdS.

One generic result:

all stable circular (if exist) geodesics in all SSS spacetimes contain 3 infinite sequences of conjugate points to any initial point and only one sequence has clear geometric meaning.

Globally maximal timelike curves

The search for globally longest curve from p to p_1 is different conceptually and in practice.

 Ω — the space of *all* future directed timelike piecewise smooth curves from *p* to *p*₁. Each curve λ has length $s(\lambda) > 0$.

 $\lambda \in \Omega$ is globally maximal if it is the longest curve in $\Omega \Leftrightarrow$ its length is equal to the Lorentzian distance function

$$s(\lambda) = d(p, p_1).$$

The maximal curve is always a geodesic (non unique).

In globally hyperbolic spacetimes for any pair $p \prec \prec p_1$ there is globally maximal geodesic $\gamma \in \Omega$.

In the global problem the conjugate point is replaced by a *cut point*.

A geodesic γ , $0 \le s < a$, is usually not globally maximal beyond some its segment.

Let

$$s_0 \equiv \sup\{s \in [0,a) : d(\gamma(0),\gamma(s)) = s\},\$$

if $s_0 < a \Rightarrow \gamma(s_0)$ — future timelike cut point of $\gamma(0)$ on γ . s_0 — the length of the longest maximal segment of γ . This means that on the segment:

- from $\gamma(0)$ to all $\gamma(s)$, $s < s_0$ γ is the unique globally maximal,
- from $\gamma(0)$ to $\gamma(s_0) \gamma$ is globally maximal (not unique),
- from $\gamma(0)$ to $\gamma(s_1)$, $s_1 > s_0$ there are curves longer than γ .

The future cut point of $p = \gamma(0)$ along γ comes *no later* than the first future conjugate point to p. A geodesic may contain only the cut point and no conjugate points.

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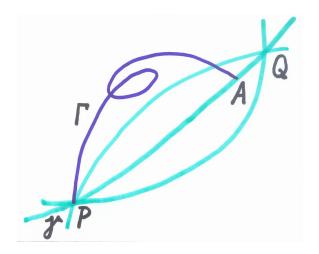


Figure: Q — nearest conjugate point of P on γ , A — the future cut point of P, Γ and γ — two longest timelike geodesics from P to A.

GLOBAL LORENTZIAN GEOMETRY

Our current knowledge:

There is a number of 'existence theorems' on maximal timelike geodesics in various spacetimes valid if some global conditions are satisfied.

They are not 'constructive': do not indicate a computationally effective procedure for finding out the interesting object \Rightarrow do *not* provide analytic tools to establish if the given geodesic is globally maximal.

This is consequence of the *nonlocal* nature of the globally maximal curve — it cannot be identified by a *local* tool such as a differential equation. Accessible tools: high symmetry and use of the Gaussian normal geodesic coordinate frame.