Invariant Solutions of the Wheeler-DeWitt equation in Hybrid Gravity

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Field equations

The action of the hybrid metric-Palatini gravity is

$$S = \frac{1}{2\kappa^2} \int d^4x \sqrt{-g} [R + f(\mathcal{R})] + S_m, \tag{1}$$

where R is a metric Ricci curvature scalar and $f(\mathcal{R})$ is a function of the Palatini curvature scalar which is constructed by an independent torsionless connection $\hat{\Gamma}$. Here $\mathcal{R}=g^{\mu\nu}\mathcal{R}_{\mu\nu}(\hat{\Gamma})$.

The modified field equations are

$$R_{\mu\nu} - \frac{1}{2}Rg_{\mu\nu} + f'(\mathcal{R})\mathcal{R}_{\mu\nu} - \frac{1}{2}f(\mathcal{R})g_{\mu\nu} = \kappa^2 T_{\mu\nu},$$
 (2)

The trace of (2) is called the hybrid structural equation. The Palatini curvature \mathcal{R} can be expressed algebraically in terms of X, assuming that $f(\mathcal{R})$ has analytic solutions:

$$f'(\mathcal{R})\mathcal{R} - 2f(\mathcal{R}) = \kappa^2 T + R \equiv X.$$
 (3)

Field equations

The action (1) is equivalent to the following one with the scalar field

$$S = \frac{1}{2\kappa^2} \int d^4x \sqrt{-g} [R + \phi \mathcal{R} - V(\phi)]. \tag{4}$$

where $\phi \equiv f'(\mathcal{R})$ and $V(\phi) = \mathcal{R}f'(\mathcal{R}) - f(\mathcal{R})$. Furthermore, for the two tensors $R_{\mu\nu}$ and $\mathcal{R}_{\mu\nu}$ it holds that¹

$$\mathcal{R}_{\mu\nu} = R_{\mu\nu} + \frac{3}{2} \frac{f(\mathcal{R})_{,\mu} f(\mathcal{R})_{,\nu}}{f^2(\mathcal{R})} - \frac{f(\mathcal{R})_{;\mu\nu}}{f(\mathcal{R})} - \frac{1}{2} \frac{\Box f(\mathcal{R})}{f(\mathcal{R})} g_{\mu\nu}$$

¹S. Capozziello, T. Harko, T.S. Koivisto, F.S.N. Lobo, G.J Olmo, JCAP 04 (2013) 011 (arXiv:1209.2895)

Field equations

By using the last relation the action (4) becomes

$$S = \frac{1}{2\kappa^2} \int d^4x \sqrt{-g} [(1+\phi)R + \frac{3}{2\phi} \partial^{\mu}\phi \partial_{\mu}\phi - V(\phi)]$$
 (5)

which is the action of a non minimally coupled scalar field. For the FRW spatially flat spacetime and for empty space ($T_{\mu\nu}=0$) the Lagrangian of the field equations is

$$\mathcal{L} = 6a\dot{a}^2(1+\phi) + 6a^2\dot{a}\dot{\phi} + \frac{3}{2\phi}a^3\dot{\phi}^2 + a^3V(\phi). \tag{6}$$

where the field equations are the Hamiltonian of (6) and the Euler-Lagrange equations with respect to the variables $x^i=(a,\phi)$. The WDW equation which is a quantization of the Hamiltonian has a form $\Delta \Psi - a^3 V(\phi) \Psi = 0$, where Δ is the Laplace operator.

Point Symmetries

Let $H\left(x^{i},u,u_{,i},u_{,ij}\right)=0$ be a PDE while $X=\xi^{i}\left(x^{j},u\right)\partial_{i}+\eta\left(x^{j},u\right)\partial_{u}$ is the generator of an infinitesimal transformation in the space $\left\{x^{i},u\right\}$. We shall say that X is a Lie point symmetry of H if there exists a function λ , such that $X^{[2]}H=\lambda H$, modH=0 where $X^{[2]}$ is the second prologation of X.

In order to select potentials $V\left(\phi\right)$ where the WDW eq. admits Lie point symmetries we will follow the geometric approach of A. Paliathanasis, M. Tsamparlis, IJGMMP (2014) 14500376, (arXiv:1312.3942) where Lie point symmetries are related to the conformal algebra of a minisuperspace.

Results

ullet If $V\left(\phi
ight)=V_{0}\left(\sqrt{\phi}+V_{1}
ight)^{4}$, the generic symmetry vector is

$$X_{\Psi}=-rac{1}{2}\partial_a+rac{\phi+V_1\sqrt{\phi}}{a}\partial_\phi+c_2\Psi\partial_\Psi$$
 $f(\mathcal{R})=rac{\mathcal{R}^2}{4V_0} \ \ ext{for} \ \ V_1=0.$

• If $V\left(\phi\right) = V_0 \left(1+\phi\right)^2 \exp\left(\frac{6}{V_1}\arctan\sqrt{\phi}\right)$, the generic symmetric vectors are $X_1 = \partial_u$, $X_{\Psi} = \Psi \partial_{\Psi}$ $X_2 = e^{-\frac{3\nu}{V_1}} \left[\cos\left(V_C u\right)\cos\left(3 v\right) + \sin\left(V_C u\right)\sin\left(3 v\right)\right] \partial_u + e^{-\frac{3\nu}{V_1}} \left[\begin{array}{c} \left(V_1\cos\left(3 v\right) - \sin\left(3 v\right)\right)\cos\left(V_C u\right) + \\ + \left(\cos\left(3 v\right) + V_1\sin\left(3 v\right)\right)\sin\left(V_C u\right) \end{array} \right] \partial_v$ $X_3 = e^{-\frac{3\nu}{V_1}} \left[\cos\left(V_C u\right)\sin\left(3 v\right) + \sin\left(V_C u\right)\cos\left(3 v\right)\right] \partial_u + e^{-\frac{3\nu}{V_1}} \left[\begin{array}{c} \left(\cos\left(3 v\right) + V_1\sin\left(3 v\right)\right)\cos\left(V_C u\right) + \\ + \left(\sin\left(3 v\right) - V_1\cos\left(3 v\right)\right)\sin\left(V_C\right) \end{array} \right] \partial_v$

Power law potential

Invariant solution

For the power law potential $V_0\left(\sqrt{\phi}+V_1\right)^4$ there exist a coordinate system $(a,\phi) \to (x,y)$ where the WDW becomes

$$\Psi_{,xx} + \Psi_{,yy} - 2V_0y^4\Psi = 0.$$

Thus by applying the zero order invariants of X_{Ψ} , $\{y, Ye^{\mu x}\}$ we can find the invariant solution

$$\Psi(x,y) = \sum_{\mu} \left[y_1 e^{\mu x + w(y)} + y_2 e^{\mu x - w(y)} \right]$$
 (7)

where $w\left(y\right)=\frac{\sqrt{2}}{2}\int\sqrt{(2ar{V}_{0}y^{4}-\mu^{2})}dy$.

Power law potential $V\left(\phi ight)=V_{0}\left(\sqrt{\phi}+V_{1} ight)^{4}$

Lie point symmetries of the WDW eq. can be used in order to construct Noether point symmetries for a Lagrangian of field equations (see arXiv:1312.3942). We have the following results.

- If $V_1=0$, then $a\left(t\right)=a_0\sqrt{t}$, i.e. the radiation solution
- If $V_1 \neq 0$, then $a(\tau) = a_0(\tau \tau_0) + a_1 a_2 \frac{1}{\tau \tau_0}$, where $dt = a(\tau) d\tau$. However, if $a_0 = 0$ the Friedmann eq. H^2 can be written

$$\frac{H^2}{H_0^2} = \Omega_{0,r} a^{-4} + \Omega_{0,m} a^{-3} + \Omega_{0,k} a^{-2} + \Omega_{0,f} a^{-1} + \Omega_{0,\Lambda}$$

where
$$\Omega_{0,i}=\Omega_{0,i}\left(a_1
ight)$$
, $i=\left\{r,\mathit{m},\mathit{k},\mathit{f},\Lambda\right\}$ and $a_2=\frac{(|a_1|+1)^2}{H_0}$.

Similarly, we can find an exact classical solution for the second potential.

Classical solution

Conclusion

- The Lie point symmetries of the WDW eq. in Hybrid Gravity were studied.
- The Lie invariants were used in order to find exact solution of the WDW and to solve analytically the modified field equations.
- It is of interest that in the case of the power law potential $V\left(\phi\right)=V_0\left(\sqrt{\phi}+V_1\right)^4$ the Friedmann equation H^2 is a fourth order polynomial with non vanishing coefficients; that is, every power law term of $\sqrt{\phi}$ in the potential produces a corresponding fluid in the model.

$$\frac{H^2}{H_0^2} = \Omega_{0,r}a^{-4} + \Omega_{0,m}a^{-3} + \Omega_{0,k}a^{-2} + \Omega_{0,f}a^{-1} + \Omega_{0,\Lambda}$$



Thank you!