

Response of gravitational wave detectors in metric theories of gravity

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- 1 **Detector response**
- 2 **Sensitivity: monochromatic signal**
- 3 **Sensitivity: stochastic signal**
- 4 **Examples: LISA, eLISA, BBO**

plane monochromatic wave with polarization π moving along Ω

$$e^{i\omega(t-\Omega\cdot\mathbf{x})}\epsilon^\pi$$

$\pi = sl, st$ (scalar), vx, vy (vector), tp, tc (tensor)

ϵ^π are the polarization tensors with components:

$$\begin{aligned}\tilde{\epsilon}^{sl} &= \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 1 \end{pmatrix}, & \tilde{\epsilon}^{st} &= \begin{pmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 0 \end{pmatrix}, \\ \tilde{\epsilon}^{vx} &= \begin{pmatrix} 0 & 0 & 1 \\ 0 & 0 & 0 \\ 1 & 0 & 0 \end{pmatrix}, & \tilde{\epsilon}^{vy} &= \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & 1 & 0 \end{pmatrix}, \\ \tilde{\epsilon}^{tp} &= \begin{pmatrix} 1 & 0 & 0 \\ 0 & -1 & 0 \\ 0 & 0 & 0 \end{pmatrix}, & \tilde{\epsilon}^{tc} &= \begin{pmatrix} 0 & 1 & 0 \\ 1 & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix}\end{aligned}$$

Detector response

$\Delta T_{aba}(t)$ – time of flight: $a \rightarrow b \rightarrow a$

$$\frac{\Delta T_{aba}(t)}{T} = \underbrace{\mathcal{T}(\omega L, \boldsymbol{\Omega} \cdot \mathbf{n}_{ab}) F^\pi(\mathbf{n}_{ab})}_{H(\omega L, \boldsymbol{\Omega}) - \text{detector response}} e^{i\omega(t - \boldsymbol{\Omega} \cdot \mathbf{x}_a)}$$

$$\mathcal{T}(x, c) = \frac{1}{2} \left[\text{sinc}\left[\frac{x}{2}(1+c)\right] e^{-\frac{ix}{2}(1+c)} + \text{sinc}\left[\frac{x}{2}(1-c)\right] e^{-\frac{ix}{2}(3+c)} \right] \quad \text{frequency transfer}$$

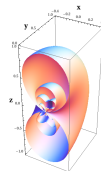
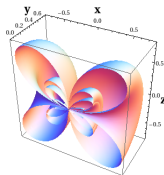
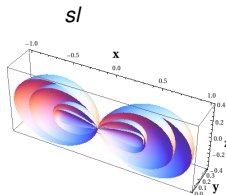
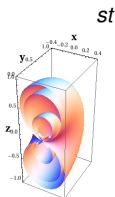
$$F^\pi(\mathbf{n}) = \frac{1}{2} \mathbf{n} \otimes \mathbf{n} : \boldsymbol{\epsilon}^\pi \quad \text{angular pattern}$$

Detector response = frequency transfer $\times$ angular pattern

Frequency transfer



Frequency transfer $\times$ angular pattern (detector arm along $\mathbf{x}$)



Monochromatic gravitational wave

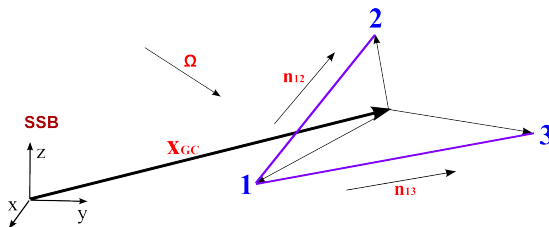
- tensor, vector

$$\mathbf{h}(\mathbf{x}, t) = h_0^{\pi_1} \cos[\omega(t - \boldsymbol{\Omega} \cdot \mathbf{x}) + \phi_0] \boldsymbol{\epsilon}^{\pi_1} + h_0^{\pi_2} \sin[\omega(t - \boldsymbol{\Omega} \cdot \mathbf{x}) + \phi_0] \boldsymbol{\epsilon}^{\pi_2}$$

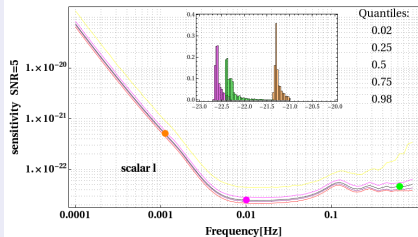
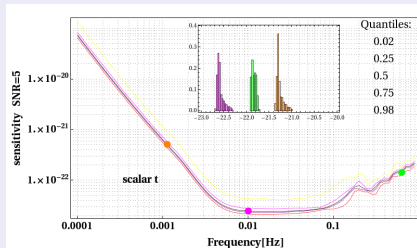
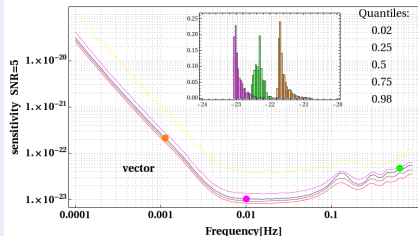
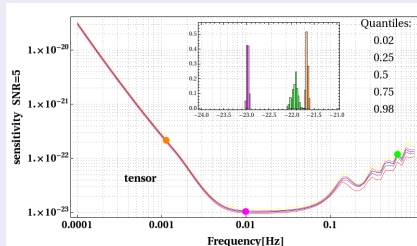
- scalar t, scalar l

$$\mathbf{h}(\mathbf{x}, t) = h_0^\pi \cos[\omega(t - \boldsymbol{\Omega} \cdot \mathbf{x}) + \phi_0] \boldsymbol{\epsilon}^\pi$$

Triangular configuration:

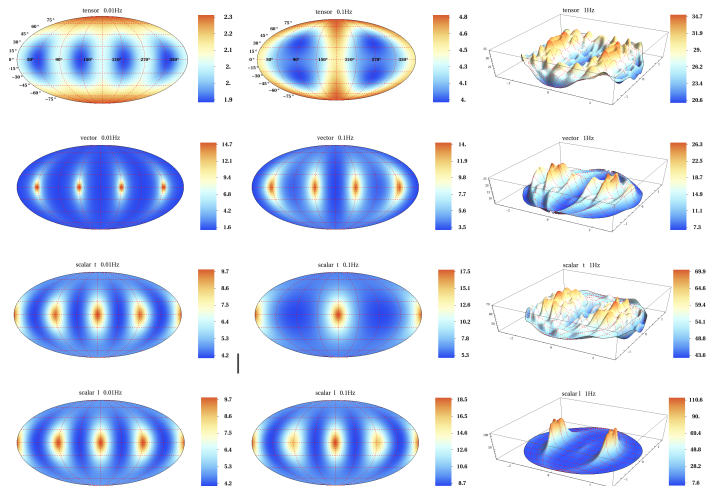


Monochromatic gravitational wave



eLISA, X ; uniform distribution of sky location and $\langle$ orientations $\rangle$.

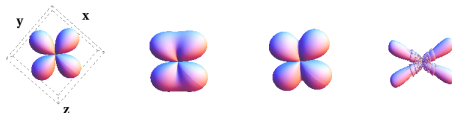
Sensitivity: sky map; SNR = 5; 10mHz



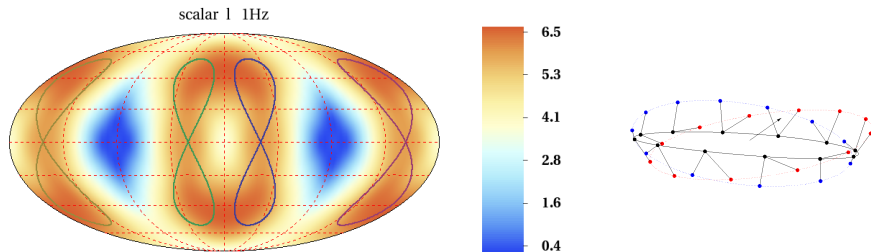
eLISA, X ; sensitivity $[10^{-23}]$; $\langle$ uniform distribution of orientations $\rangle$.

SNR for sl mode

sl response for eLISA arms:



SNR ($h = 10^{-22}$); uniform distribution of orientations.



Triangular configuration - stochastic signal

- stochastic background: self- and cross- correlations

$$\langle s_I(t) s_J(t) \rangle = \frac{1}{2} \int_{-\infty}^{\infty} \int_{S^2} S_h(\omega, \Omega) \mathcal{F}_{IJ}(\omega, \Omega) \times \text{filter} \times \frac{d\Omega}{4\pi} d\omega,$$

$S_h(\omega, \Omega)$ - possibly anisotropic background

- antenna pattern

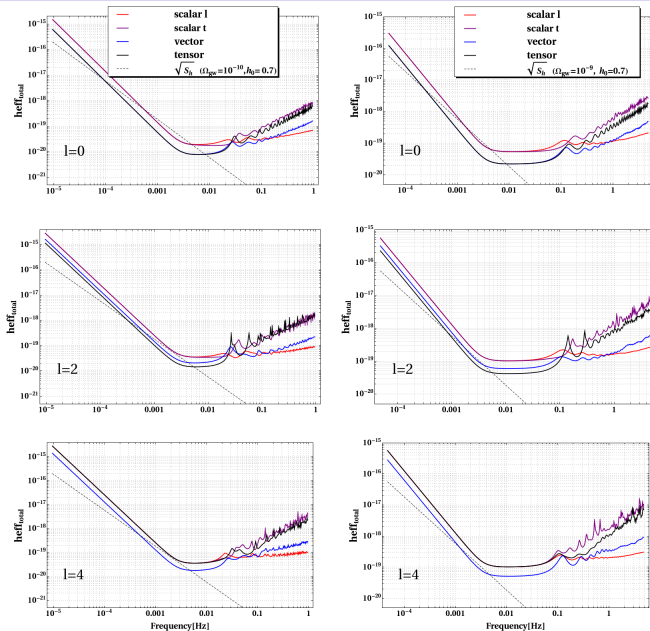
$$\mathcal{F}_{IJ}(\omega, \Omega) = e^{i\omega \Omega \cdot (\mathbf{x}_J - \mathbf{x}_I)} [\mathcal{T}_I(\omega L, \Omega \cdot \mathbf{n}_I) F^\pi(\mathbf{n}_I)] [\mathcal{T}_J(\omega L, \Omega \cdot \mathbf{n}_J) F^\pi(\mathbf{n}_J)]^*$$

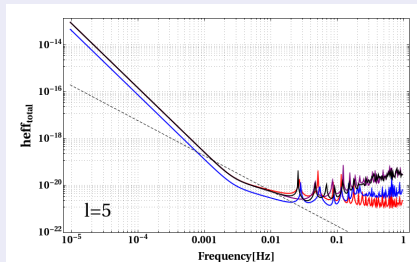
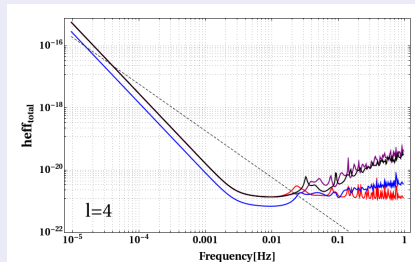
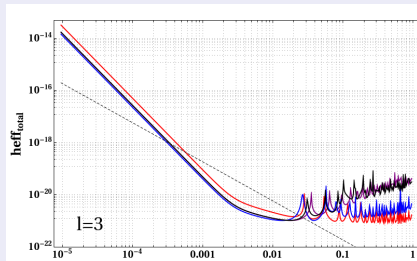
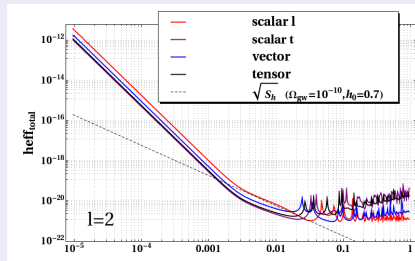
- spherical harmonic decomposition of $\mathcal{F}_{IJ}$: $a_{lm}(\omega L) = \int_{S^2} Y_{lm}^*(\Omega) \mathcal{F}(\omega L, \Omega) \frac{d\Omega}{4\pi}$

angular power of the antenna pattern:
$$\sigma_I(\omega L) = \sqrt{\frac{1}{2I+1} \sum_{m=-I}^I |a_{Im}(\omega L)|^2}$$

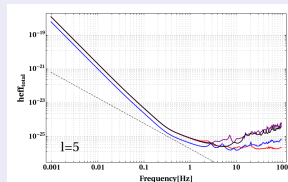
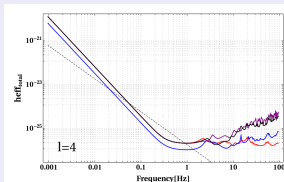
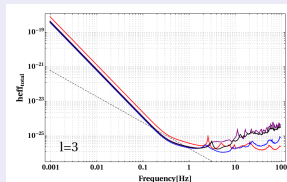
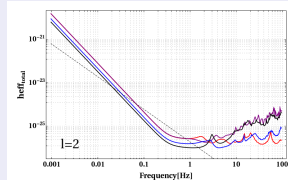
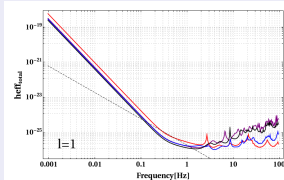
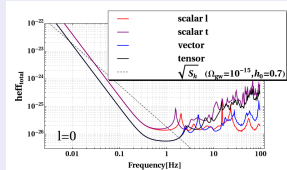
- sensitivity

self-correlations: $h_{eff} = (4\pi)^{1/4} \sqrt{\frac{S_{II}(f)}{\sigma_I^2(f)}}$, cross-correlations: $h_{eff} = \left(\frac{4\pi}{2I+1}\right)^{1/4} \left[\frac{S_{IJ}(f) S_{JI}(f)}{\sigma_I^2(f)} \right]^{1/4}$





cross-corr: **AE**, $T = 1\text{yr}$



cross-corr: **AA** + **AE** + $\dots$ + **TT**, $T = 1\text{yr}$

References:

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