An approach to initial conditions in general relativity

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Tafel J. 2014, An approach to initial conditions in general relativity, sent to Class. Quantum Grav., arXiv: 1405.1202

Tafel J. and Jóźwikowski M. 2014, New solutions of initial conditions in general relativity, *Class. Quantum Grav.* **31** 115001, arXiv:1312.7819

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Notation: $T^{ij} = K^{ij} - Hg^{ij}$



$$g'_{ij} = \psi^4 g_{ij}, \quad K'_{ij} = \psi^{-2} K_{ij} + \frac{1}{3} H' g_{ij}, \quad K^i_{i} = 0$$

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(the momentum constraint with H=0). Thus, it makes sense to solve the momentum constraint first, provided that H can be set to zero.

The momentum constraint in the Gauss coordinates

Initial metric

$$g = g_{ab}dx^a dx^b + d\varphi^2$$
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$$(|\tilde{g}|T^{33})_{,3} = |\tilde{g}|(g_{ab,3}T^{ab} - T_{a|a}^{3a}),$$

where

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Problems with integration if H=0 or ∂_{φ} is a symmetry.

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$$\theta^a = dx^a$$
, $\theta^3 = d\varphi + \beta_a dx^a$
 $e_a = \partial_a - \beta_a \partial_3$, $e_3 = \partial_3$.

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$$\begin{split} \theta^{a} &= dx^{a} \;,\;\; \theta^{3} = d\varphi + \beta_{a}dx^{a} \\ e_{a} &= \partial_{a} - \beta_{a}\partial_{3} \;,\;\; e_{3} = \partial_{3} \;. \end{split}$$

The momentum constraint

$$(|\tilde{g}|T_3^3)_{,3} - \frac{1}{2}|\tilde{g}|g_{ab,3}T^{ab} = 2|\tilde{g}|\beta_{a,3}T_3^a - \frac{1}{\alpha}e_a(\alpha|\tilde{g}|T_3^a)$$

$$(\alpha T_{a}{}^{b})_{|b} - |\tilde{g}|^{-1} (\alpha |\tilde{g}| T_{a}{}^{b} \beta_{b})_{,3} = (e_{a} \alpha - \alpha \beta_{a,3}) T_{3}{}^{3} - \frac{1}{2} \alpha (g_{bc,3} T^{bc}) \beta_{a} + \alpha \lambda \eta_{ab} T_{3}{}^{b} - |\tilde{g}|^{-1} (\alpha |\tilde{g}| T_{a}{}^{3})_{,3},$$

where η^{ab} is the Levi-Civita tensor and $\lambda=\eta^{ab}e_a\beta_b$.

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Functions equivalent to $K_{ij} - \frac{1}{3}Hg_{ij}$:

$$U = \frac{1}{2}\alpha(T_{11} - T_{22}) + i\alpha T_{12} , \quad V = \alpha(T_{13} + iT_{23}) , \quad W = T_3^3 + \frac{2}{3}H .$$

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A function of V and the metric (no U and W)

$$E = \rho \alpha^{-1} \text{Re}(2\beta_{,3}V - \partial V) + \frac{2}{3}\rho^{3}H_{,3} \; , \; \; \partial = \partial_{\xi} - \beta\partial_{3} \; .$$

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A function of V, W and the metric (no U)

$$F = \frac{1}{2}\rho^2\alpha^{-2}\bar{\partial}(\alpha^3W) - \frac{3}{2}\rho^2\alpha W\bar{\beta}_{,3} - iIm(\partial\bar{\beta})V - (\alpha^{-2}\rho^2V)_{,3} + \frac{2}{3}\rho^2\alpha\bar{\partial}H \; .$$

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For analytic data

$$U = \bar{\chi}_{,3} \left(\int_{\xi_0}^{\xi} d\xi' \; \frac{F}{\bar{\chi}_{,3}} + h(\bar{\xi}, \bar{\chi}) \right) ,$$

where $\beta=\bar{\chi}_{,\xi}/\bar{\chi}_{,3}$ and the integrand is considered as a function of $\xi',\ \bar{\xi}$ and $\bar{\chi}.$

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In the nonanalytic case with eta=0

$$U = -rac{1}{4\pi}\int d^2x'rac{F(x'^a,arphi)}{(ar{\xi'}-ar{\xi})} + h(ar{\xi},arphi) \ .$$

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Comments

If $U=\hat{U}_{,3}$ and $F=\hat{F}_{,3}$ then

$$\partial \hat{U} = \hat{F}$$
.

Let the CR structure corresponding to ∂ be realizable: $\partial \bar{\chi} = 0$ and $\chi \neq f(\xi)$. Equivalently

$$\beta = \frac{\bar{\chi}_{,\xi}}{\bar{\chi}_{,3}} \ .$$

Then

$$\hat{U} = \int_{\xi_0}^{\xi} d\xi' \; \hat{F}(\xi', ar{\xi}, ar{\chi}) + \hat{h}(ar{\xi}, ar{\chi})$$

and taking the φ -derivative yields U.

Formula for $\beta=0$ follows from the fundamental solution of the Laplace equation.



Axially symmetric data (paper with Jóźwikowski)

Then

$$\partial V = 0$$
.

Equivalently, in terms of real fields and coordinates,

$$T_3^{\ a} = \alpha^{-1} \eta^{ab} \omega_{,b} \ .$$

Other components of T^{ij} are given in quadrature. In generic case they can be found explicitly but then H=0 is a differential equation.

Families of data generalizing the Kerr data can be constructed (up to a knowledge of ψ).

There exists a class of data (defined in quadrature) which satisfy all constraints but they are not asymptotically flat.

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Theorem (D. Maxwell)

Let (S,g) be a complete Riemannian manifold without boundary and let (g,K) be asymptotically flat data of class $W_{\delta}^{k,p}$.

There exists a unique solution of the Lichnerowicz equation such that $\psi>0$ and $(\psi-1)\in W^{k,p}_\delta$ if and only if

$$||f||_{L^{6}}^{2} \leq \lambda (\int_{S} (8|\nabla f|^{2} + Rf^{2}) dv_{g}), \quad \lambda > 0$$

for all $f \in C_c^{\infty}$.

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Full system of constraints for $H' \neq const$

$$\triangle \psi = \frac{1}{8} R \psi - \frac{1}{8} K_{ij} K^{ij} \psi^{-7} + \frac{1}{8} H^{2} \psi^{5}$$

$$(\rho^3 W)_{,3} = \frac{2}{3} \psi^6 \rho^3 H'_{,3} + E_0$$

$$U_{,\xi} - (\beta U)_{,3} = \frac{2}{3} \psi^6 \rho^2 \alpha \bar{\partial} H' + F_0$$

where E_0 and F_0 denote E and F with H=0.

J. Tafel (Institute of Theoretical Physics An approach to initial conditions in gener

If
$$ho_{,3}=0$$
 then
$$({T_3}^3)_{,3}=\rho^{-3}E_0$$
 defines ${T_3}^3$ for given $V=\alpha(T_{13}+iT_{23}).$

If $ho_{,3}=0$ then

$$(T_3^3)_{,3} = \rho^{-3} E_0$$

defines T_3^3 for given $V = \alpha(T_{13} + iT_{23})$.

Calculate H from the Hamiltonian constraint

$$R - \, \tilde{T}_{ab} \, \tilde{T}^{ab} - 2 \, T_{a3} \, T^{a3} - 2 (\, T_3^{\,\, 3})^2 - \, T_3^{\,\, 3} H = 0 \,\, . \label{eq:R_ab}$$

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Substitute H and T_3^3 into the equation for U. If $\beta=0$ it reads

$$U_{,\xi}-\rho^2(fU\bar{U})_{,\bar{\xi}}=h$$
.

Functions f and h are known and coordinate φ appears as a parameter.



Horizons

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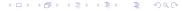
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Problems with the existence theorems unless K(n, n) = 0 and with an extension of initial data through the boundary (some solutions are constructed in the paper with Jóźwikowski).



An approach without boundary

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Initial metric induced by the Kerr solution on t = const

$$g = \rho^2 \Delta^{-1} dr^2 + \rho^2 d\theta^2 + \rho^{-2} \Sigma^2 \sin^2 \theta d\varphi^2$$

where

$$\rho^2 = r^2 + a^2 \cos^2 \theta \; , \; \; \Delta = r^2 - 2Mr + a^2 \; , \; \Sigma^2 = (r^2 + a^2)^2 - a^2 \Delta \sin^2 \theta \; .$$

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Nonvanishing components of T^{ij} :

$$T_3^a = \alpha^{-1} \eta^{ab} \omega_{,b}$$
,

$$\omega = 4aM\rho^{-2}[2(r^2 + a^2) + (r^2 - a^2)\sin^2\theta]\cos\theta.$$

Conformally flat complex coordinate: $\xi = (\tilde{r} + i\theta)/2$, where

$$r = M + \sqrt{M^2 - a^2} \cosh \tilde{r} , \ \tilde{r} \in [-\infty, \infty]$$

 $(\tilde{r} = 0 \text{ is the external Kerr horizon}).$



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- $T^{ij}\tilde{r}_{,i}\tilde{r}_{,j}=0$

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Solving the momentum constraint with respect to U and W under these assumptions is practically impossible because of the condition $T^{ij}n_in_j=0$. Let us try to solve them with respect to V.

Let

$$g = \rho^2 d\xi d\bar{\xi} + \alpha^2 d\varphi^2$$
, $\rho_{,3} = \alpha_{,3} = H = 0$

and functions ρ , α and T^{ij} be even functions of $\tilde{r}=\frac{1}{2}Re\xi$.

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$$2ReU - \alpha \rho^2 T_3^3 = 0 \quad at \quad \tilde{r} = 0$$

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$$Re((\rho^{-2}\alpha^2 U_{,\xi})_{,\xi}) = \frac{1}{2}(\alpha^3 T_3^3)_{,\xi\bar{\xi}} + \rho^2 \alpha^{-2}(\alpha^3 T_3^3)_{,33}$$

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then momentum constraint can be solved with respect to V. If there exists a unique solution ψ of the Lichnerowicz equation then the initial data admit a MOTS at $\tilde{r}=0$.

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- For analytic data or reduced number of free functions the momentum constraint can be solved in quadrature and the conformal method can be applied if H=0.
- The algebraic method of solving the Hamiltonian constraint and the construction of a generalized Kerr data with MOTS are (perhaps) worth studying.
- More work needed to include multiple horizons.